

Europa Initiative team meeting
Optimisation of the European orbits for JEM

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with the help of :
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- ① Interplanetary cruise
 - Transfers using swing-by's
 - Direct transfer to Jupiter

- ② From JOI to EOI

- ③ Delivery & Relay phase
 - Purpose of those orbits
 - Frozen orbits
 - Halo orbits

- ④ Science Orbit
 - Connecting halo orbits to science orbits
 - Results

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Transfers using swing-by's

| Launch Date | Flyby Path | TOF to JOI (yrs.) | C_3 (km ² /s ²) |
|--------------------|------------|-------------------|--|
| 25 Mar 2020 | VEE | 6.03 | 15.6 |
| 27 May 2021 | VEE | 6.87 | 14.5 |
| 21 Nov 2021 | VEE | 6.37 | 15.0 |
| 15 May 2022 | EVEE | 7.22 | 10.2 |
| 23 May 2023 | VEE | 6.18 | 16.4 |
| 03 Sep 2024 | VEE | 6.71 | 13.8 |
| 01 Aug 2026 | VEE | 6.94 | 10.0 |
| 21 Jul 2026 | VEE | 6.15 | 15.2 |

Figure 1: Short list of interplanetary trajectories (from ES 2012 Report by NASA)

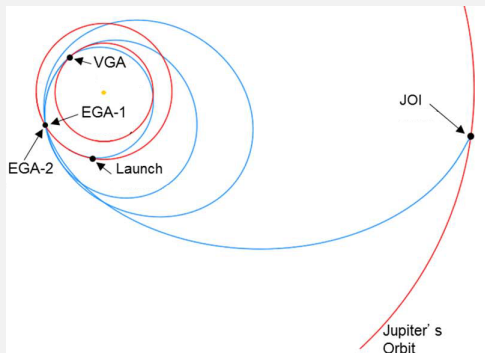


Figure 2: Example of VEEGA trajectory

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Direct transfer to Jupiter

Hohmann transfer :
 $\Delta V_1 = 8.80 \text{ km/s}$
 $\Delta V_2 = 5.64 \text{ km/s}$
TOF = 2.73 years

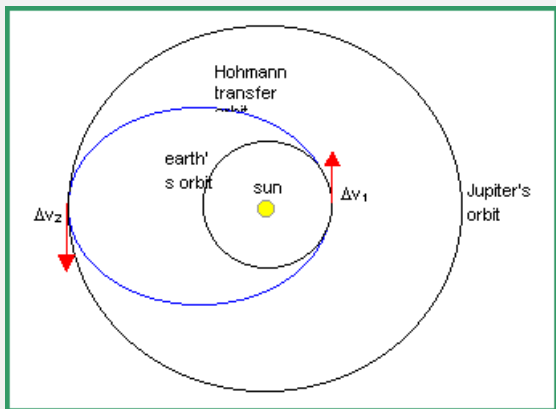


Figure 3: Hohmann transfer

Jovian tour

Table B.2.6-4. Trade-offs between Flight-time, deterministic ΔV , and TID (Si behind 100 mil Al, spherical shell) for various types of tours as compared to the concept baseline tour

| Flight Time (delta yrs) | ΔV (delta km/s) | JOI-to-EOI, inclusive | |
|----------------------------|----------------------------|-----------------------|--|
| | | TID (delta Mrad) | Type of Tour |
| 0 | > 5.5 | ~0 | No tour, direct insertion to Europa Orbit from interplanetary trajectory |
| 0.25 | 4 | ~0 | Callisto gravity assists and v-infinity leveraging |
| 0.5 | 3 | ~0 | Further Callisto gravity assists and v-infinity leveraging |
| 1 | 2.5 | 0.1–0.5 | Callisto and Ganymede gravity assists (no endgame) |
| 1.5 | 1.5 | 0.8–1.2 | Callisto, Ganymede, and Europa gravity assists (4:3, 6:5 endgame) |
| 2.5 | 1.3 | 1.7 | Callisto, Ganymede, Europa and Io gravity assists |

Figure 4: Trade-offs between Flight-time, ΔV , and TID (from ES 2012 Report by NASA)

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Landing site

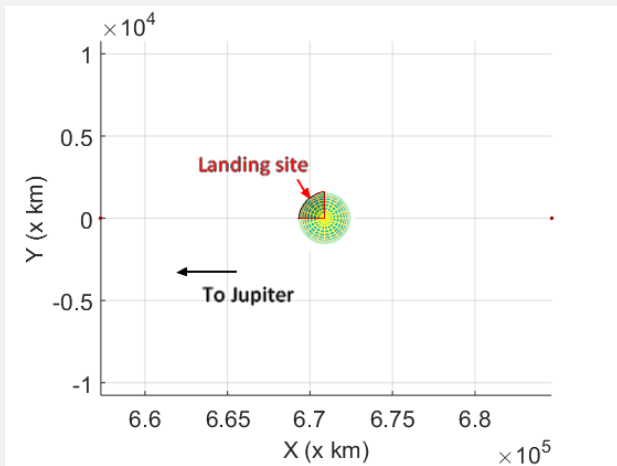


Figure 5: Restricted area for landing site in Jupiter-Europa rotating frame (Projection on Europa Orbital plane)

Delivery : The carrier spacecraft must be within line of sight of the lander stack from deployment to landing in order to provide communications coverage for critical events associated with landing.

Relay : ensure communication with lander around 60% of time

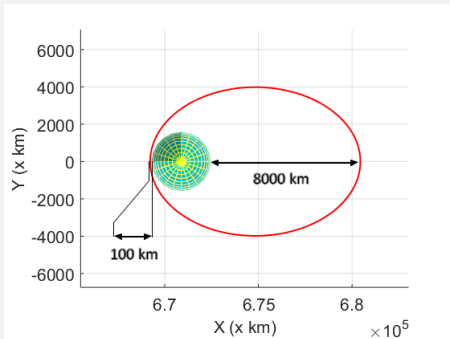


Figure 6: Class keplerian orbit

Pb : Periastre is circulating around Europa (explain how)

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Frozen orbits

fixed periaspe

low-inclined and stable : expensive to go on a near-polar orbit

R.Russell is investigating them

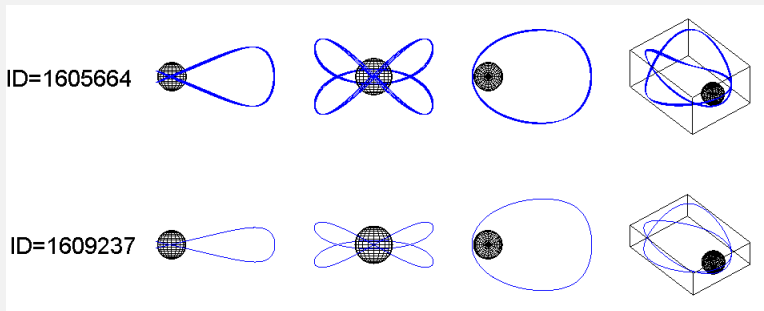


Figure 7: Stable orbits around Europa in the rotating frame (Ryan P. Russell)

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Halo orbits

- Unstable periodic orbits around L1 & L2
- Symmetric about XZ-plane
- Two families : Northern & Southern
- Easy to leave
- Little maneuvers (few m/s for delivery and relay phase) required to stay on orbit
- Maneuvers computed from Earth

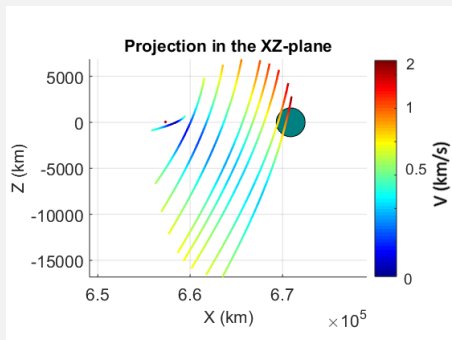


Figure 8: JEL1 Halo orbits

Example of halo orbit

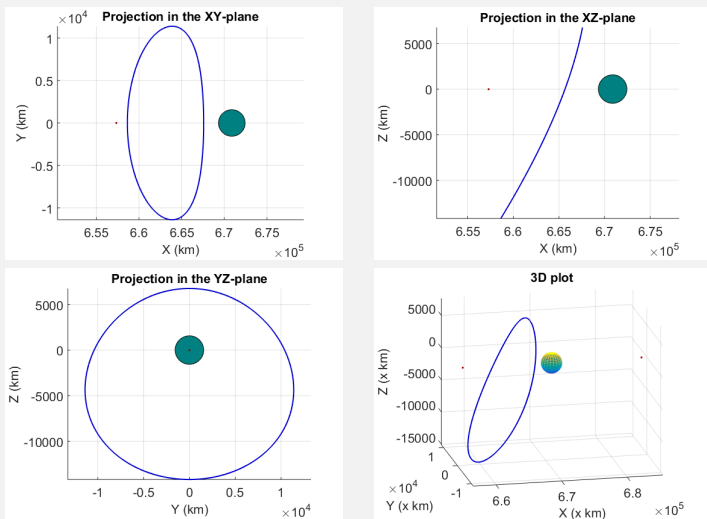


Figure 9: Projection and 3D-view of a JEL1 halo orbit ($C = 3.0013$) in the rotating frame

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Connecting halo orbits to science orbits

| | Min. | Max. |
|-------------------|------|----------|
| altitude (km) | 100 | 200 |
| inclination (deg) | 80 | 90 |
| duration (days) | 30 | ∞ |

Table 1: Science orbit characteristics

Description of the algorithm :

- Choose a halo orbit
- Choose a location on halo orbit
- Tangent burn or small burn in the unstable direction
- Search for extrema location
- Second burn to circularize around Europa
- Compute circular orbit until crash on Europa

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Example of transfers ($C_{halo} = 3.0013$)

| ΔV_1 (m/s) | ΔV_2 (m/s) | ΔV_{tot} (m/s) | TOF (days) | altitude (km) | inclination ($^\circ$) |
|--------------------|--------------------|------------------------|------------|---------------|--------------------------|
| 8.0 | 487.5 | 495.5 | 1.31-30.0 | 190 | 83.8 |
| 216 | 449.6 | 665.6 | 0.51 | 109.7 | 86.1 |
| 48 | 480.7 | 528.7 | 0.93 | 146.5 | 80.9 |
| 40 | 481.4 | 521.4 | 1.34 | 186.9 | 88.3 |

Table 2: Transfers with a tangent burn

| ΔV_1 (m/s) | ΔV_2 (m/s) | ΔV_{tot} (m/s) | TOF (days) | altitude (km) | inclination ($^\circ$) |
|--------------------|--------------------|------------------------|------------|---------------|--------------------------|
| 0.8 | 490.1 | 490.9 | 2.27 | 194.1 | 80.5 |
| 0.5 | 502.9 | 503.3 | 2.91 | 127.5 | 82.1 |
| 2.8 | 489.4 | 492.1 | 1.83 | 198.3 | 80.6 |
| 5.6 | 492.4 | 497.9 | 1.64 | 181.7 | 80.3 |

Table 3: Transfers with small burn in the unstable direction

Example of transfers ($C_{halo} = 3.0013$)

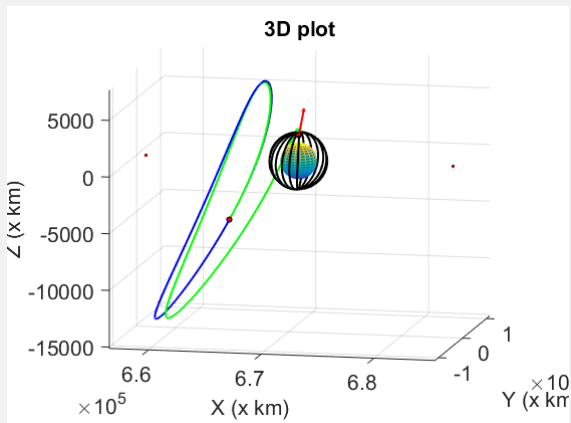


Figure 10: Transfer from a JEL2 Halo orbit ($C = 3.0013$) to a science orbit ($\Delta V_{tot} = 495.5\text{m/s}$, TOF = 1.31d, alt=190km, inc = 83.8°)